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# Some Characterizations of Space-Like Rectifying Curves in the Minkowski Space-Time 

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# Some Characterizations of Space-Like Rectifying Curves in the Minkowski Space-Time 

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#### Abstract

In this work, a space-like rectifying curve with space-like principal normal in the Minkowski space-time $E_{1}^{4}$ is defined as a curve whose position vector always lies in orthogonal complement $N^{\perp}$ of its principal normal vector field $N$. Also, we characterized such curves in terms of their curvature functions and we obtained the necessary and sficient conditions for such curve to be a rectifying curve.


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## I. INTRODUCTION

Lorentzian geometry helps to bridge the gab between modern differential geometry and the mathematical physics of general relativity by giving an invariant treatment of Lorentzian geometry. The fact that relativity theory is expressed in terms of Lorentzian geometry is attractive for geometers, who can penetrate surprising quickly into cosmology (redshift, expanding universe and big bang) and a topic no less interesting geometrically, the gravitation of a single star (perihelion procession, bending of light and black holes) [18]. Despite its long history, the theory of curve is still one of the most important interesting topics in a differential geometry and its is being study by many mathematicians until now, see for example $[1,2,3,4,5,16,19,21,24]$.

In the Euclidean space $E^{3}$, rectifying curves are introduced by Chen in [7] as space curves whose position vector always lies in its rectifying plane, spanned by the tangent and the binormal vector fields $\vec{T}$ and $\vec{B}$ of the curve. Therefore, the position vector $\vec{\alpha}$ of a rectifying curve satisfies the equation

$$
\vec{\alpha}(s)=\lambda(s) \vec{T}(s)+\mu(s) \vec{B}(s),
$$

for some differentiable functions $\lambda$ and $\mu$ in arclength function $s$. The Euclidean rectifying curves are studied in [7, 8]. In particular, it is shown in [8] that there exists a simple relationship between the rectifying curves and the centrodes, which play some important roles in mechanics, kinematics as well as in differential geometry in defining the curves of constant precession. The rectifying curves are also studied in [8] as the extremal curves. In the Minkowski 3 -space $E_{1}^{3}$, the rectifying curves are investigated in [10]. The rectifying curves are also studied in [11] as the centrodes and extremal curves. In the Euclidean 4 -space $E^{4}$, the rectifying curves are investigated in [9].

In analogy with the rectifying curve, the curve whose position vector always lies in its normal plane spanned by the principal normal and the binormal vector fields $\vec{N}$ and $\vec{B}$ of the curve is called normal curve in Euclidean 3space $E^{3}$ and it is well known that normal curves are spherical curves in $E^{3}$ [8]. Similar definition and characterizations of space-like, time-like (and also null) and dual time-like normal curves are given in references [10, 11, 17]. The space-like normal curve in Minkowski 4 -space $E_{1}^{4}$ is defined in [14] as a curve whose position vector always lies in the orthogonal complement $\vec{T}^{\perp}$ of its tangent vector field $\vec{T}$ which is given by

$$
\vec{T}^{\perp}=\left\{\vec{W} \in E_{1}^{4} \mid g(\vec{W}, \vec{T})=0\right\}
$$

In [6], Camci and others have shown that a space - like curve lies in pseudohyperbolic space $H_{0}^{3}$ iff the following equation holds

$$
\vec{\alpha}-m=-\left(1 / k_{1}\right) \vec{N}-\left(1 / k_{2}\right)\left(1 / k_{1}\right)^{\prime} \overrightarrow{B_{1}}+\left(1 / k_{3}\right)\left[k_{2} / k_{1}+\left(\left(1 / k_{2}\right)\left(1 / k_{1}\right)^{\prime}\right)^{\prime}\right] \overrightarrow{B_{2}},
$$

Where $m$ is constant, $k_{1}, k_{2}$ and $k_{3}$ are the first, the second and the third curvatures of the curve $\alpha$, respectively. By using the definition of space-like normal curves in Minkowski 4-space $E_{1}^{4}$ and the last equality, it follows that every space-like curve lying in pseudohy-perbolic space $H_{0}^{3}$ is a normal curve in Minkowski 4-space.

In this paper, in analogy with the Minkowski 3-dimensional case, we define the rectifying curve in the Minkowski 4 -space $E_{1}^{4}$ as a curve whose position vector always lies in the orthogonal complement $\vec{N} \perp$ of its principal normal vector field $\vec{N}$ Consequently, $\vec{N}^{\perp}$ is given by

$$
\vec{N}^{\perp}=\left\{\vec{W} \in E_{1}^{4} \mid g(\vec{W}, \vec{N})=0\right\}
$$

[^0]Where $g(.,$.$) denotes the standard pseudo scalar product in E_{1}^{4}$. Hence $\vec{N}^{\perp}$ is a 3 -dimensional subspace of $E_{1}^{4}$, spanned by the tangent, the first binormal and the second binormal vector fields $\vec{T}, \vec{B}_{1}$ and $\vec{B}_{2}$ respectively. Therefore, the position vector with respect to some chosen origin, of a space-like rectifying curve $\vec{\alpha}$ in Minkowski space-time $E_{1}^{4}$, satisfies the equation

$$
\begin{equation*}
\vec{\alpha}(s)=\lambda(s) \vec{T}(s)+\mu(s) \vec{B}_{1}(s)+\nu(s) \vec{B}_{2}, \tag{1}
\end{equation*}
$$

for some differentiable functions $\lambda(s), \mu(s)$ and $\nu(s)$ in arclength function $s$. Next, characterize space-like rectifying curves in terms of their curvature functions $\kappa_{1}(s), \kappa_{2}(s)$ and $\kappa_{3}(s)$ and give the necessary and the sufficient conditions for arbitrary curve in $E_{1}^{4}$ to be a rectifying. Moreover, we obtain an explicit equation of a space-like rectifying curve in $E_{1}^{4}$ and give the relation between rectifying and normal space-like curves in $E_{1}^{4}$.

## iI. Preliminaries

In this section, we prepare basic notations on Minkowski space-time $E_{1}^{4}$. Let $\vec{\alpha}: I \subset R \rightarrow E_{1}^{4}$ be arbitrary curve in the Minkowski space - time $E_{1}^{4}$. Recall that the curve $\vec{\alpha}$ is said to be unit speed (or parameterized by arclength function $s$ ) if $g\left(\vec{\alpha}^{\prime}, \vec{\alpha}^{\prime}\right)= \pm 1$, where $g(.,$.$) denotes the standard pseudo scalar product in E_{1}^{4}$ given by

$$
g(\vec{v}, \vec{w})=-v_{1} w_{1}+v_{2} w_{2}+v_{3} w_{3}+v_{4} w_{4},
$$

for each $\vec{v}=\left(v_{1}, v_{2}, v_{3}, v_{4}\right) \in E_{1}^{4}$ and $\vec{w}=\left(w_{1}, w_{2}, w_{3}, w_{4}\right) \in E_{1}^{4}$. An arbitrary vector $\vec{v} \in E_{1}^{4}$ can have one of three Lorentzian causal characters; it can be space - like if $g(\vec{v}, \vec{v})>0$ or $\vec{v}=0$, time - like if and null (light-like) if $g(\vec{v}, \vec{v})=0$ and $\vec{v} \neq 0$. Similarly, an arbitrary curve $\vec{\alpha}=\vec{\alpha}(s)$ can locally be space-like, timelike or null (light-like), if all of its velocity vectors $\vec{\alpha}^{\prime}(s)$ are respectively space-like, time-like or null (light-like). Also recall that the pseudo-norm of an arbitrary vector $\vec{v} \in E_{1}^{4}$ is given by $\|\vec{v}\|=\sqrt{|g(\vec{v}, \vec{v})|}$. Therefore $\vec{v}$ is a unit is a unit vector if $g(\vec{v}, \vec{v})= \pm 1$. The velocity of the curve $\vec{\alpha}(s)$ is given by $\left\|\vec{\alpha}^{\prime}(s)\right\|$. Next, vectors $\vec{v}, \vec{w}$ in $E_{1}^{4}$ are said to be orthogonal if $g(\vec{v}, \vec{w})=0$.
Denote by $\left\{\vec{T}(s), \vec{N}(s), \vec{B}_{1}(s), \vec{B}_{2}(s)\right\}$ the moving Frenet frame along the curve $\vec{\alpha}(s)$ in the space $E_{1}^{4}$, where $\vec{T}(s)$, $\vec{N}(s), \vec{B}_{1}(s)$ and $\vec{B}_{2}(s)$ are the tangent, principal normal, the first binormal and second binormal fields, respectively. For an arbitrary space - like curve $\vec{\alpha}(s)$ with space - like principal normal $\vec{N}$ in the space $E_{1}^{4}$, the following Frenet formula are given in [23, 20, 6, 15, 22, 25]:

$$
\left[\begin{array}{c}
\vec{T}^{\prime}  \tag{2}\\
\vec{N}^{\prime} \\
\vec{B}_{1}^{\prime} \\
\vec{B}_{2}^{\prime}
\end{array}\right]=\left[\begin{array}{cccc}
0 & \kappa_{1} & 0 & 0 \\
-\kappa_{1} & 0 & \kappa_{2} & 0 \\
0 & -\varepsilon \kappa_{2} & 0 & \kappa_{3} \\
0 & 0 & \kappa_{3} & 0
\end{array}\right]\left[\begin{array}{c}
\vec{T} \\
\vec{N} \\
\vec{B}_{1} \\
\vec{B}_{2}
\end{array}\right]
$$

Where

$$
\begin{equation*}
g\left(\vec{B}_{1}, \vec{B}_{1}\right)=-g\left(\vec{B}_{2}, \vec{B}_{2}\right)=\varepsilon= \pm 1, \quad g(\vec{T}, \vec{T})=g(\vec{N}, \vec{N})=1 . \tag{3}
\end{equation*}
$$

Recall the functions $\kappa_{1}(s), \kappa_{2}(s)$ and $\kappa_{3}(s)$ are called respectively, the first, the second and the third curvatures of space-like curve $\vec{\alpha}(s)$. Here, $\varepsilon$ determines the kind of space - like curve $\alpha(s)$. If $\varepsilon=1$, then $\alpha(s)$ is a sp ace like curve with space - like first binormal $\vec{B}_{1}$ and time - like second binormal $B_{2}$. If $\varepsilon=-1$, then $\alpha(s)$ is a space-like curve with time - like first binormal $\vec{B}_{1}$ and space - like second binormal $B_{2}$. If $\kappa_{3}(s) \neq 0$ for each $s \in I \subset R$, the curve $\vec{\alpha}$ lies fully in $E_{1}^{4}$. Recall that the pseudohyperbolic space $H_{0}^{3}(1)$ in $E_{1}^{4}$, centered at the origin, is the hyperquadric defined by

$$
H_{0}^{3}(1)=\left\{\vec{X} \in E_{1}^{4} \mid g(\vec{X}, \vec{X})=-1\right\} .
$$

Recently, Yilmaz et al. [27, 26] defined a vector product in Minkowski spacetime $\mathbf{E}_{1}^{4}$ as follows:
Definition 2.1. Let $a=\left(a_{1}, a_{2}, a_{3}, a_{4}\right), b=\left(b_{1}, b_{2}, b_{3}, b_{4}\right)$ and $c=\left(c_{1}, c_{2}, c_{3}, c_{4}\right)$ be vectors in $\boldsymbol{E}_{1}^{4}$. The vector product in Minkowski space -time $E_{1}^{4}$ is defined by the determinant

$$
a \wedge b \wedge c=-\left[\begin{array}{cccc}
-e_{1} & e_{2} & e_{3} & e_{4}  \tag{4}\\
a_{1} & a_{2} & a_{3} & a_{4} \\
b_{1} & b_{2} & b_{3} & b_{4} \\
c_{1} & c_{2} & c_{3} & c_{4}
\end{array}\right]
$$

where $e_{1}, e_{2}, e_{3}$ and $e_{4}$ are mutually orthogonal vectors (coordinate direction vectors) satistying equations

$$
e_{1} \wedge e_{2} \wedge e_{3}=e_{4}, e_{2} \wedge e_{3} \wedge e_{4}=e_{1}, e_{3} \wedge e_{4} \wedge e_{1}=e_{2}, e_{4} \wedge e_{1} \wedge e_{2}=-e_{3}
$$

Lemma 2.2. Let $a=\left(a_{1}, a_{2}, a_{3}, a_{4}\right), b=\left(b_{1}, b_{2}, b_{3}, b_{4}\right)$ and $c=\left(c_{1}, c_{2}, c_{3}, c_{4}\right)$ be vectors in $\boldsymbol{E}_{1}^{4}$. From the definition of vector product, there is a property in Minkowski space - time $E_{1}^{4}$ as the following:

$$
\begin{equation*}
g(a \wedge b \wedge c, a)=g(a \wedge b \wedge c, b)=g(a \wedge b \wedge c, c)=0 \tag{5}
\end{equation*}
$$

The proof of above lemma is elementary.

## ili. Some Characterizations of Rectifying Curves In $E_{1}^{4}$

In this section, we firstly characterize the space-like rectifying curves with space-like principal normal in Minkowski space - time in terms of their curvatures. Let $\vec{\alpha}=\vec{\alpha}(s)$ be a unit speed space - like rectifying curve in $E_{1}^{4}$, with non-zero curvatures $\kappa_{1}(s), \kappa_{2}(s)$ and $\kappa_{3}(s)$ By definition, the position vector of the curve $\vec{\alpha}$ sacrifies the equation (1), for some differentiable functions $\lambda(s), \mu(s)$ and $\nu(s)$. Differentiating the equation (1) with respect to $s$ and using the Frenet equations (2), we obtain

$$
\begin{equation*}
\vec{T}=\lambda^{\prime} \vec{T}+\left(\lambda \kappa_{1}-\varepsilon \mu \kappa_{2}\right) \vec{N}+\left(\mu^{\prime}+\nu \kappa_{3}\right) \vec{B}_{1}+\left(\nu^{\prime}+\mu \kappa_{3}\right) \vec{B}_{2} . \tag{6}
\end{equation*}
$$

It follows that

$$
\begin{align*}
& \lambda^{\prime}=1, \\
& \lambda \kappa_{1}-\varepsilon \mu \kappa_{2}=0, \\
& \mu^{\prime}+\nu \kappa_{3}=0,  \tag{7}\\
& \nu^{\prime}+\mu \kappa_{3}=0,
\end{align*}
$$

and therefore

$$
\begin{align*}
& \lambda=s+c \\
& \mu=\varepsilon \frac{\kappa_{1}(s)(s+c)}{\kappa_{2}},  \tag{8}\\
& \nu=-\varepsilon \frac{\kappa_{1}(s) \kappa_{2}(s)+(s+c)\left(\kappa_{1}^{\prime}(s) \kappa_{2}(s)-\kappa_{1}(s) \kappa_{2}^{\prime}(s)\right)}{\kappa_{2}^{2}(s) \kappa_{3}(s)},
\end{align*}
$$

where $c \in R$. In this way the functions $\lambda(s), \mu(s)$ and $\nu(s)$ are expressed in terms of the curvature functions $\kappa_{1}(s), \kappa_{2}(s)$ and $\kappa_{3}(s)$ of the curve $\alpha(s)$. Moreover, by using the last equation in (7) and relation (8) we easily find that the curvatures $\kappa_{1}(s), \kappa_{2}(s)$ and $\kappa_{3}(s)$ satisfy the equation

$$
\begin{equation*}
\frac{\kappa_{1}(s) \kappa_{3}(s)(s+c)}{\kappa_{2}(s)}-\left[\frac{\kappa_{1}(s) \kappa_{2}(s)+(s+c)\left(\kappa_{1}^{\prime}(s) \kappa_{2}(s)-\kappa_{1}(s) \kappa_{2}^{\prime}(s)\right)}{\kappa_{2}^{2}(s) \kappa_{3}(s)}\right]^{\prime}=0, \quad c \in R . \tag{9}
\end{equation*}
$$

The condition (9) can be written as:

$$
\begin{equation*}
\frac{\kappa_{1}(s)(s+c)}{\kappa_{2}(s)}-\frac{1}{\kappa_{3}(s)} \frac{d}{d s}\left[\frac{1}{\kappa_{3}(s)} \frac{d}{d s}\left(\frac{\kappa_{1}(s)(s+c)}{\kappa_{2}(s)}\right)\right]=0 . \tag{10}
\end{equation*}
$$

If we change the variable $s$ by the variable $t$ as the following

$$
\frac{d}{d t}=\frac{1}{\kappa_{3}(s)} \frac{d}{d s} \text { or } t=\int_{0}^{s} \kappa_{3}(s) d s
$$

the equation (10) takes the following form

$$
\begin{equation*}
\frac{\kappa_{1}(s)(s+c)}{\kappa_{2}(s)}-\frac{d^{2}}{d t^{2}}\left[\frac{\kappa_{1}(s)(s+c)}{\kappa_{2}(s)}\right]=0 . \tag{11}
\end{equation*}
$$

General solution of this equation is

$$
\begin{equation*}
\frac{\kappa_{1}(s)(s+c)}{\kappa_{2}(s)}=\varepsilon\left(A \cosh \int_{0}^{s} \kappa_{3}(s) d s+B \sinh \int_{0}^{s} \kappa_{3}(s) d s\right) \tag{12}
\end{equation*}
$$

where $A$ and $B$ are arbitrary constants. Then from (8) we have

$$
\begin{align*}
& \lambda(s)=s+c \\
& \mu(s)=A \cosh \int_{0}^{s} \kappa_{3}(s) d s+B \sinh \int_{0}^{s} \kappa_{3}(s) d s  \tag{13}\\
& \nu(s)=-\left(A \sinh \int_{0}^{s} \kappa_{3}(s) d s+B \cosh \int_{0}^{s} \kappa_{3}(s) d s\right) .
\end{align*}
$$

Conversely, assume that the curvatures $\kappa_{1}(s), \kappa_{2}(s)$ and $\kappa_{3}(s)$ of an arbitrary unit speed space - like curve in $E_{1}^{4}$, satisfy the equation (12). Let us consider the vector $\vec{X} \in E_{1}^{4}$ given by

$$
\begin{align*}
\vec{X}(s) & =\vec{\alpha}(s)-(s+c) \vec{T}(s)-\left(A \cosh \int_{0}^{s} \kappa_{3}(s) d s+B \sinh \int_{0}^{s} \kappa_{3}(s) d s\right) \vec{B}_{1}(s)  \tag{14}\\
& +\left(A \sinh \int_{0}^{s} \kappa_{3}(s) d s+B \cosh \int_{0}^{s} \kappa_{3}(s) d s\right) \vec{B}_{2}(s)
\end{align*}
$$

By using the relations (2) and (12), we easily find $\vec{X}^{\prime}(s)=0$, which means that $\vec{X}$ is a constant vector. This implies that $\alpha(s)$ is congruent to a rectifying curve. In this way, the following theorem is proved.
Theorem 3.1. Let $\vec{\alpha}(s)$ be unit speed space-like curve with space-like principal normal in $E_{1}^{4}$ and with non -zero curvatures $\kappa_{1}(s), \kappa_{2}(s)$ and $\kappa_{3}(s)$.Then $\vec{\alpha}(s)$ is congruent to a space -like rectifying curve if and only if

$$
\frac{\kappa_{1}(s)(s+c)}{\kappa_{2}(s)}=\varepsilon\left(A \cosh \int_{0}^{s} \kappa_{3}(s) d s+B \sinh \int_{0}^{s} \kappa_{3}(s) d s\right) .
$$

In particular, assume that all curvature functions $\kappa_{1}(s), \kappa_{2}(s)$, and $\kappa_{3}(s)$ of space - like rectifying curve, $\vec{\alpha}$ in $E_{1}^{4}$ are constant and different from zero. Then equation (9) easily implies a contradiction. Hence we obtain the following theorem.
Theorem 3.2. There are no space-like rectifying curves with space - like principal normal Iying in $E_{1}^{4}$, with non-zero constant curvatures $\kappa_{1}(s), \kappa_{2}(s)$ and $\kappa_{3}(s)$.
In the next theorem, we give the necessary and the suficient conditions for the space - like curve $\alpha(s)$ in $E_{1}^{4}$ to be a rectifying curve.
Theorem 3.3. Let $\alpha(s)$ be unit speed space-like rectifying curve with space-like prini-pal normal in $E_{1}^{4}$, with nonzero curvatures $\kappa_{1}(s), \kappa_{2}(s)$ and $\kappa_{3}(s)$. Then the following statements hold:
(i) The distance function $\rho(s)=\|\vec{\alpha}(s)\|$ satisfies $\rho^{2}(s)=s^{2}+c_{1} s+c_{2}, c_{1} \in R$ and $c_{2} \in R_{0}$.
(ii) The tangential component of the position vector of the space-like rectitying curve is given by $g(\vec{\alpha}(s), \vec{T}(s))=$ $s+c, c \in R$.
(iii) The normal component $\vec{\alpha}^{N}(s)$ of the position vector of the space - like rectifying curve has constant length and the distance function $\rho(s)$ is non-constant.
(iv) The first binormal component and the second binormal component of the position vector of the space-like rectifying curve are respectively given by

$$
\begin{align*}
& g\left(\vec{\alpha}(s), \vec{B}_{1}(s)\right)=\varepsilon\left(A \cosh \int_{0}^{s} \kappa_{3}(s) d s+B \sinh \int_{0}^{s} \kappa_{3}(s) d s\right) \\
& g\left(\vec{\alpha}(s), \vec{B}_{2}(s)\right)=\varepsilon\left(A \sinh \int_{0}^{s} \kappa_{3}(s) d s-B \cosh \int_{0}^{s} \kappa_{3}(s) d s\right) . \tag{15}
\end{align*}
$$

Conversely, if $\vec{\alpha}(s)$ is a unit speed curve in $E_{1}^{4}$ with non - zero curvatures $\kappa_{1}(s), \kappa_{2}(s)$ and $\kappa_{3}(s)$ and one of the statements (i), (ii), (iii) or (iv) holds, then $\vec{\alpha}(s)$ is a space - like rectifying curve.
Proof. Let us first suppose that $\vec{\alpha}(s)$ is a unit speed space - like rectifying curve in $E_{1}^{4}$ with non - zero curvatures $\kappa_{1}(s), \kappa_{2}(s)$ and $\kappa_{3}(s)$. The position vector of the curve $\vec{\alpha}(s)$ satisfies the equation (1), where the functions $\lambda(s), \mu(s)$ and $\nu(s)$ satisfy relation (13). From relation (1) and (13) we have

$$
\begin{align*}
g(\vec{\alpha}, \vec{\alpha}) & =\lambda^{2}+\varepsilon\left(\mu^{2}(s)-\nu^{2}(s)\right),  \tag{16}\\
& =(s+c)^{2}+\varepsilon\left(A^{2}-B^{2}\right)
\end{align*}
$$

Therefore, $\rho^{2}(s)=s^{2}+c_{1} s+c_{2}, c_{1} \in R$ and $c_{2} \in R_{0}$, which proves statement (i).
But using the relations (1) and (8) we easily get $g(\vec{\alpha}(s), \vec{T}(s))=s+c, c \in R$, so the statement (ii) is proved.

Note that the position vector of an arbitrary curve $\vec{\alpha}(s)$ in $E_{1}^{4}$ can be decomposed as $\vec{\alpha}(s)=$ $m(s) \vec{T}(s)+\vec{\alpha}^{N}(s)$, where $m(s)$ is arbitrary differentiable function and $\vec{\alpha}^{N}(s)$ is the normal component of the position vector. If $\vec{\alpha}(s)$ is a space - like rectifying curve, relation (1) implies $\vec{\alpha}^{N}(s)=\mu(s) \vec{B}_{1}(s)+\nu(s) \vec{B}_{2}(s)$ and therefore $g\left(\vec{\alpha}^{N}(s), \vec{\alpha}^{N}(s)\right)=\varepsilon\left(\mu^{2}(s)-\nu^{2}(s)\right)$.Moreover, by using (13), we find $\left\|\vec{\alpha}^{N}(s)\right\|=\varepsilon\left(A^{2}-B^{2}\right)=a$, $a \in R$. By statement (i), $\rho(s)$ is non -constant function, which proves statement (iii).
Finally, using (1), (3) and (13) we easily obtain (15), which proves statement (iv).
Conversely, assume that statement (i) holds. Then $g(\vec{\alpha}(s), \vec{\alpha}(s))=s^{2}+c_{1} s+c_{2}, c_{1} \in R, c_{2} \in R$
By differentiating the previous equation two times with respect to $s$ and using (2), we obtain $g(\vec{\alpha}(s), \vec{N}(s))=0$, which implies that $\vec{\alpha}$ is a space - like rectifying curve.
If statement (ii) holds, in a similar way it follows that $\vec{\alpha}$ is a space - like rectifying curve.
If statement (iii) holds, let us put $\vec{\alpha}(s)=m(s) \vec{T}(s)+\vec{\alpha}^{N}(s)$, where $m(s)$ is arbitrary differentiable function. Then

$$
\begin{equation*}
g\left(\vec{\alpha}^{N}(s), \vec{\alpha}^{N}(s)\right)=g(\vec{\alpha}(s), \vec{\alpha}(s))-2 g(\vec{\alpha}(s), \vec{T}(s)) m(s)+m^{2}(s) . \tag{17}
\end{equation*}
$$

Since $g(\vec{\alpha}(s), \vec{T}(s))=m(s)$, it follows that

$$
\begin{equation*}
g\left(\vec{\alpha}^{N}(s), \vec{\alpha}^{N}(s)\right)=g(\vec{\alpha}(s), \vec{\alpha}(s))-g(\vec{\alpha}(s), \vec{T}(s))^{2} \tag{18}
\end{equation*}
$$

where $g(\vec{\alpha}(s), \vec{\alpha}(s))=\rho^{2}(s) \neq$ constant. Defferentiating the previous equation with respect to $s$ and using (2), we find

$$
\begin{equation*}
\kappa_{1}(s) g(\vec{\alpha}(s), \vec{T}(s)) g(\vec{\alpha}(s), \vec{N}(s))=0 . \tag{19}
\end{equation*}
$$

It follows that $g(\vec{\alpha}(s), \vec{N}(s))=0$ and hence the space - like curve $\vec{\alpha}$ is a rectifying.
If the statement (iv) holds, by taking the derivative of the equation

$$
\begin{equation*}
g\left(\vec{\alpha}(s), \vec{B}_{1}(s)\right)=\varepsilon\left(A \cosh \int_{0}^{s} \kappa_{3}(s) d s+B \sinh \int_{0}^{s} \kappa_{3}(s) d s\right) \tag{20}
\end{equation*}
$$

with respect to $s$ and using (2), we obtain

$$
\begin{equation*}
-\varepsilon \kappa_{2}(s) g(\vec{\alpha}(s), \vec{N}(s))+\kappa_{3} g\left(\vec{\alpha}(s), \vec{B}_{2}(s)\right)=\varepsilon \kappa_{3}\left(A \sinh \int_{0}^{s} \kappa_{3}(s) d s+B \cosh \int_{0}^{s} \kappa_{3}(s) d s\right) \tag{21}
\end{equation*}
$$

By using (15), the last equation becomes $g(\vec{\alpha}(s), \vec{N}(s))=0$, which means that $\vec{\alpha}$ is a space - like rectifying curve. This proves the theorem.
In the next theorem, we find the parametric equation of a rectifying curve.
Theorem 3.4. Let $\alpha: I \subset R \rightarrow E_{1}^{4}$ be a space - like curve with space -like principal normal in $E_{1}^{4}$ given by $\vec{\alpha}$ $(t)=\rho(t) \vec{y}(t)$ where $\rho(t)$ is an arbitrary positive function and $\vec{y}(t)$ is a unit speed space - like curve lying in pseudohyperbolic space $H_{0}^{3}(1)$. Then $\vec{\alpha}$ is a space -like rectifying curve if and only if

$$
\begin{equation*}
\rho(t)=\frac{a}{\cosh \left(t+t_{0}\right)}, \quad a \in R_{0}, \quad t_{0} \in R \tag{22}
\end{equation*}
$$

Proof. Let $\vec{\alpha}$ be a curve in $E_{1}^{4}$ given by

$$
\begin{equation*}
\vec{\alpha}(t)=\rho(t) \vec{y}(t) \tag{23}
\end{equation*}
$$

where $\rho(t)$ is arbitrary positive function and $\vec{y}(t)$ is a unit speed space - like curve in the pseudohyperbolic space $H_{0}^{3}(1)$.. By taking the derivative of the previous equation with respect to $t$, we get

$$
\begin{equation*}
\vec{\alpha}^{\prime}(t)=\rho^{\prime}(t) \vec{y}(t)+\rho(t) \vec{y}^{\prime}(t) \tag{24}
\end{equation*}
$$

Hence the unit tangent vector of $\vec{\alpha}$ is given by

$$
\begin{equation*}
\vec{T}=\frac{\rho^{\prime}(t)}{v(t)} \vec{y}(t)+\frac{\rho(t)}{v(t)} \vec{y}^{\prime}(t) \tag{25}
\end{equation*}
$$

where $v(t)=\left\|\vec{\alpha}^{\prime}(t)\right\|$ is the speed of $\vec{\alpha}$. Differentiating the equation (25) with respect to $t$, we find

$$
\begin{equation*}
\vec{T}^{\prime}=\left(\frac{\rho^{\prime}}{v}\right)^{\prime} \vec{y}+\left(\frac{2 \rho^{\prime}}{v}-\frac{\rho \rho^{\prime}\left(\rho+\rho^{\prime \prime}\right)}{v^{3}}\right) \vec{y}^{\prime}+\left(\frac{\rho}{v}\right) \vec{y}^{\prime \prime} . \tag{26}
\end{equation*}
$$

Let $\vec{Y}$ be the unit vector field in $E_{1}^{4}$ satisfying the equations $g(\vec{Y}, \vec{y})=g\left(\vec{Y}, \vec{y}^{\prime}\right)=0$. Then $\left\{\vec{y}, \vec{y}^{\prime}, \vec{Y}, \vec{y} \wedge \vec{y}^{\prime}\right.$ $\wedge \vec{Y}\}$ is orthonormal frame in $E_{1}^{4}$. Therefore decomposition of $\vec{y}^{\prime \prime}$ with respect the frame $\left\{\vec{y}, \vec{y}^{\prime}, \vec{Y}, \vec{y} \wedge \vec{y}^{\prime} \wedge \vec{Y}\right\}$ reads

$$
\begin{align*}
\vec{y}^{\prime \prime}= & g\left(\vec{y}^{\prime \prime}, \vec{y}\right) g(\vec{y}, \vec{y}) \vec{y}+g\left(\vec{y}^{\prime \prime}, \vec{y}^{\prime}\right) g\left(\vec{y}^{\prime}, \vec{y}^{\prime}\right) \vec{y}^{\prime}+g\left(\vec{y}^{\prime \prime}, \vec{Y}\right) g(\vec{Y}, \vec{Y}) \vec{Y} \\
& +g\left(\vec{y}^{\prime \prime}, \vec{y} \wedge \vec{y}^{\prime} \wedge \vec{Y}\right) g\left(\vec{y} \wedge \vec{y}^{\prime} \wedge \vec{Y}, \vec{y} \wedge \vec{y}^{\prime} \wedge \vec{Y}\right) \vec{y} \wedge \vec{y}^{\prime} \wedge \vec{Y} . \tag{27}
\end{align*}
$$

Since $g(\vec{y}, \vec{y})=-1$ and $g\left(\vec{y}^{\prime}, \vec{y}^{\prime}\right)=1$, it follows that $g\left(\vec{y}^{\prime \prime}, \vec{y}\right)=-1$ and $g\left(\vec{y}^{\prime \prime}, \vec{y}^{\prime}\right)=0$, so the equation (27) the equation (27) becomes

$$
\begin{equation*}
\vec{y}^{\prime \prime}=\vec{y}+g\left(\vec{y}^{\prime \prime}, \vec{Y}\right) \vec{Y}+g\left(\vec{y}^{\prime \prime}, \vec{y} \wedge \vec{y}^{\prime} \wedge \vec{Y}\right) \vec{y} \wedge \vec{y}^{\prime} \wedge \vec{Y} . \tag{28}
\end{equation*}
$$

Substituting (28) into (26) and applying Frenet formulas for arbitrary speed curves in $E_{1}^{4}$, we find

$$
\begin{align*}
\kappa_{1} v \vec{N} & =\left[\left(\frac{\rho^{\prime}}{v}\right)^{\prime}+\frac{\rho}{v}\right] \vec{y}+\left(\frac{2 \rho^{\prime}}{v}-\frac{\rho \rho^{\prime}\left(\rho+\rho^{\prime \prime}\right)}{v^{3}}\right) \vec{y}^{\prime}+\left(\frac{\rho}{v}\right) g\left(\vec{y}^{\prime \prime}, \vec{Y}\right) \vec{Y}  \tag{29}\\
& +\frac{g\left(\vec{y}^{\prime \prime}, \vec{y} \wedge \vec{y}^{\prime} \wedge \vec{Y}\right)}{v} \vec{\alpha} \wedge \vec{y}^{\prime} \wedge \vec{Y} .
\end{align*}
$$

Since $g(\vec{y}, \vec{y})=-1$, we have $g\left(\vec{y}, \vec{y}^{\prime}\right)=0$ and thus $g\left(\vec{\alpha}, \vec{y}^{\prime}\right)=0$. We also have $g(\vec{\alpha}, \vec{Y})=0$. By definition, $\vec{\alpha}$ is a space - like rectifying curve in $E_{1}^{4}$ if and only if $g(\vec{\alpha}, \vec{N})=0$. Therefore, after taking the scalar product of (29) with $\vec{\alpha}$, we have $g(\vec{\alpha}, \vec{N})=0$ if and only if

$$
\begin{equation*}
\left(\frac{\rho^{\prime}}{v}\right)^{\prime}+\frac{\rho}{v}=0 \tag{30}
\end{equation*}
$$

whose general solutions are $\rho(t)=\frac{a}{\cosh \left(t+t_{0}\right)}$ or $\rho(t)=\frac{a}{\sinh \left(t+t_{0}\right)}, a \in R_{0}, t \in R$. Since, $g(\vec{T}, \vec{T})=1$, it follows that $\rho(t)=\frac{a}{\cosh \left(t+t_{0}\right)}$ is the only solution. This proves the theorem.

In Theorem 3.4, since $\vec{y}(s)$ is a unit speed space - like curve in the pseudohyperbolic space $H_{0}^{3}(1), \vec{y}(s)$ is a space - like normal curve in Minkowski 4-space $E_{1}^{4}$. So, Theorem 3.4 gives the relation between space-like rectifying and space-like normal curves in Minkowski 4-space $E_{1}^{4}$. Then we can give the following corollary.
Corollary 3.5. In Minkowski 4-space $E_{1}^{4}$, the construction of the space - like rectifying curve with space-like principal normal can be made by using the space-like normal curves.
Example: Let us consider the curve

$$
\vec{\alpha}(s)=\frac{a}{\cosh \left(s+s_{0}\right)}(\sqrt{2} \cosh (s / \sqrt{3}), \sqrt{2} \sinh (s / \sqrt{3}), \sin (s / \sqrt{3}), \cos (s / \sqrt{3})),
$$

where $a \in R_{0}, s_{0} \in R$ in $E_{1}^{4}$ This curve has a form $\vec{\alpha}(s)=\rho(s) \vec{y}(s)$ where $\rho(s)=\frac{a}{\cosh \left(s+s_{0}\right)}$ and $\vec{y}(s)=(\sqrt{2} \cosh (s / \sqrt{3}), \sqrt{2} \sinh (s / \sqrt{3}), \sin (s / \sqrt{3}), \cos (s / \sqrt{3}))$ Since $g(\vec{y}(s), \vec{y}(s))=-1$ and $\left\|\vec{y}^{\prime}(s)\right\|=1$, $\vec{y}(s)$ is a unit speed space-like curve in pseudohyper-bolic space $H_{0}^{3}(1)$. According to theorem 3.4, $\vec{\alpha}$ is a space like rectifying curve lying fully in $E_{1}^{4}$.

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